Bayesian Analysis at JET and the Minerva framework

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Problems with modern large scale scientific experiments.

Drowning in data?

Solutions: Better data aquisition systems, databases etc.

Drowning in complexity.

Solutions: ?



A top-down view on our inference problem.



The Minerva analysis infrastructure seed



Minerva

• A fully modular inference infrastructure (separating diagnostic models/combinations, physics assumptions, inversions, forward models etc)

> • Based on Bayesian Graphical Models (handles modularity of complex pdf:s)

• Handles all dependency management (keeps track of every single parameter/option that can change), so user never needs to think about data flow, this also makes automatic caching possible.

• Can fully *declare* a scientific model in a complex experiment.

Written in Java

• Machine independent (used at JET, MAST, W7-X, H-1?)

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J Svensson, A Werner, Large Scale Bayesian Data Analysis for Nuclear Fusion Experiments, WISP 2007

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Probability Engineering: Bayesian Graphical Models:

- A combination of graph theory and probability theory.
- Makes it easier to handle complex probabilistic systems (like our systems!).
- Nodes represent variables: stochastic or deterministic.
- Edges represent dependencies



Generally:
$$p(n_1, n_2, n_3, ..., n_{N_T}) = \prod_{i=1}^{N_T} p(n_i \mid par(n_i))$$

Minerva Example graphs

KG1 (interferometry)



Minerva Prediction



1400 1600 1800 2000

KG10



-1.5

0 200 400 600 800 1000 1200 Sweep frequency

-> Can be used for understanding a diagnostic/ diagnostic "debugging"/diagnostic design/ experimental planning etc



Bayes intro: An inference example

Extract temperature from Doppler broadening of spectral line:



Bayes Theorem



Inference on toroidal current distribution



Maths for current tomography

Forward function: $\overline{D}^{\text{Pred}} = \overline{\overline{M}}\overline{I} + \overline{C}$

Likelihood: Gaussian noise

$$p(\overline{D}^{Mag} | \overline{I}) = \frac{1}{(2\pi)^{N_D/2} \left| \overline{\Sigma}_D \right|^{1/2}} \exp\left(-\frac{1}{2} (\overline{\overline{M}} \overline{I} + \overline{C} - \overline{D}^{Mag})^T \overline{\Sigma}_D^{-1} (\overline{\overline{M}} \overline{I} + \overline{C} - \overline{D}^{Mag})\right)$$

Prior: Multivariate Normal over free currents (prior covariance imposes e.g. smoothing between nearby beams)

$$p(\bar{I}) = \frac{1}{(2\pi)^{N_I/2} \left| \sum_{I=1}^{m_I} \right|^{1/2}} \exp(-\frac{1}{2} (\bar{I} - \bar{m}_I)^T \sum_{I=1}^{m_I} (\bar{I} - \bar{m}_I))$$

$$\Rightarrow \text{Posterior: Multivariate normal over free currents}$$

$$p(\overline{I} \mid \overline{D}^{Mag}) = \frac{1}{(2\pi)^{N_I/2} \left| \overline{\Sigma} \right|^{1/2}} \exp(-\frac{1}{2} (\overline{I} - \overline{m})^T \overline{\Sigma}^{=-1} (\overline{I} - \overline{m})) \qquad (\text{MAximum Posterior})$$

$$where \quad \overline{m} = (\overline{M}^T \overline{\Sigma}_D \overline{M} + \overline{\Sigma}_I^{-1})^{-1} \overline{M}^T \overline{\Sigma}_D (\overline{D}^{Mag} - \overline{C})$$

$$\overline{\Sigma} = (\overline{M}^T \overline{\Sigma}_D \overline{M} + \overline{\Sigma}_I^{-1})^{-1}$$

n_e with uncertainty in mapping



j_{tor} ,p,with force balance assumption



Pickups, Saddles, Flux loops

H-mode: pedestal pressure from magnetics:



n_e and T_e profiles from combined core LIDAR, edge LIDAR and interferometer









Interim summary and outlook

• Summary:

- Uncertainties on flux surfaces, combined with inversions
- Bayesian exploration of uncertainties in equilibrium inference
- Greatly improved accuracy on ne and Te profile inference at JET
- In total about 10 diagnostic systems modelled in Minerva.
- Ongoing Minerva projects:
 - ECE diagnostics (PhD of Stefan Schmuck, KTH, Sweden)
 - Soft-X and Bolometry (PhD of Dong Li, Greifswald University)
 - MAST modelling: Thomson scattering, CX, gen. force balance (with ANU, Australia).
 - Reflectometry (with Antoine Sirinelli, JET),
 - Bremsstrahlung and He-beam (with Maciej Krychowiak, IPP)

Excursion:

Inference on infinite dimensional functions

There are two quite different classes of inversion problems:

1. We have a real underlying parameterised physics-based model. Example: fitting line spectra to (Gaussian, Voigt etc) line shapes.

2. We want to find an underlying, in principle infinite-dimensional, function, and we do not really have any specific parameterisation we believe represent the underlying physics process.

Example: tomographic inversion, density profiles etc

We tend to use methods developed for (1) also for problems belonging to class (2)!

... by choosing some "flexible" parameterisation such as grids, polynomials/splines etc. then regularizing the solution. But what kind of "smoothness" does a polynomial of A given order express in comparison to a set of Gaussian basis functions or something else?

Difference between prior constraints and parametric constraints.

Parameters "freeze" in or dictates the exact form of all possible functions we could observe, at all positions.

Prior constraints, on the other hand, provide a "loose" probabilistic constraint that can be overridden by the data.

... so should we not prefer to "guide" the inference on infinite dimensional functions by prior constraints rather than parametric constraints?



Can we do non-parametric tomography?

What happens if we let the density of beams, or density of profile points -> Infinity?



..the prior mean becomes a function over R,Z and the prior covariance matrix over beams becomes a continous *covariance function* defining the prior covariance between every pair of (R,Z):

... the posterior can be calculated directly and gives another posterior mean and a posterior covariance function.

Both prior and posterior are now *Gaussian random processes* rather than pdf:s over discrete parameters.

J Svensson, Non-parametric tomography using Gaussian processes, forthcoming

Non-parametric tomography $m_{f*} = m_{p_{f*}} + K_{*L}(K_{LL} + \Sigma_y)^{-1}(y - f_L)$ $\Sigma_{f*} = K_{**} - K_{*L}(K_{LL} + \Sigma_y)^{-1}K_{L*}$

$$(K_{**})_{kl} = k(x_*^k, x_*^l) \qquad k$$

$$(K_{L*})_{kl} = \int_{los k} k(x, x_*^l) dx \qquad y$$

$$(K_{*L})_{kl} = \int_{los l} k(x_*^k, x) dx$$

$$(K_{LL})_{kl} = \int_{los l} \int_{los l} k(x, x') dx dx'$$

k – prior covariance function

 $f_{\rm L}-predicted \ observations at prior function mean$

 \boldsymbol{x}_* – points (R,Z) where we want to evaluate inferred function y – los observations

J Svensson, Non-parametric tomography using Gaussian processes, forthcoming

Example: non-parametric interferometry inversion



J Svensson, Non-parametric tomography using Gaussian processes, forthcoming

Minerva for Wendelstein 7-X



