

me





# The physics of cold to burning plasmas



#### RSPE Seminar, 18 April 2013

Acknowledgement: Australian Research Council, ANU, DIISRTE







# The physics of cold to burning plasmas



#### CONSORZIO RFX

**Ricerca Formazione Innovazione** 



Max-Planck-Institut für Plasmaphysik

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M. J. Hole,

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International collaborators

✓ RSPE collaborators



RSPE Seminar, 18 April 2013



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### The plasma state : the fourth state of matter

plasma is an ionized gas







#### A Galaxy of Fusion Reactors.

 99.9% of the visible universe is in a plasma state



Inner region of the M100 Galaxy in the Virgo Cluster, imaged with the Hubble Space Telescope Planetary Camera at full resolution.

Fusion is the process that powers the sun and the stars

with magnetic field.

Key concepts: magnetic confinement, physics models

# **Magnetic confinement**



# **Several toroidal confinement concepts**

















MAST (UK) compact

RFX-mod (Italy) self-organising

H1-MNRF (Oz) *flexible shape* 

LHD (Japan) steady-state





W-7X (Germany) steady-state, reduced chaos

# **Physics regimes and physics models**



- Dielectric tensor: often used for cold plasmas
- MHD: flowing plasma, single temperature for ions/electrons
- Particle in cell (PIC) simulation – individual particles
- Gyrokinetic simulation simulate particle distribution functions

### How is a plasma described?

• One approach: Through dielectric tensor in Maxwell's equations

$$\nabla \cdot \mathbf{E} = \rho / \varepsilon_0 \qquad \nabla \cdot \mathbf{B} = 0$$
$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \qquad \nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \varepsilon_0 \frac{\partial \mathbf{E}}{\partial t} = \mu_0 \frac{\partial \mathbf{D}}{\partial t}$$

What is dispersion relation for a cold plasma with **B**<sub>0</sub>, **E**<sub>0</sub>=0, **v**<sub>0</sub>=0?

- Gauss's law requires charge neutrality  $\rho = 0 = \sum q_s n_s$
- Solve momentum equation

$$m\dot{\mathbf{v}}_1 = q\big(\mathbf{E}_1 + \mathbf{v}_1 \times \mathbf{B}_0\big)$$

for wave field  $v_1$ . Gives  $J=n_0 q v_1$ , and construct  $\varepsilon$  s.t.  $D = \varepsilon E$ 

- Rearrange Gauss's law, Amperes law to give  $\nabla \times \nabla \times \mathbf{E} = \frac{\omega^2}{c^2} \boldsymbol{\epsilon} \cdot \mathbf{E}$
- Search for wave like solutions  $\mathbf{E} = |\mathbf{E}(\mathbf{x})| \exp[i(\mathbf{k} \cdot \mathbf{x} \omega t)]$ giving dispersion relation  $\omega(\mathbf{k})$

### or... Magnetohydrodynamics (MHD)

• Single conducting fluid:  $\mathbf{J} = n_i Z e \mathbf{v}_i - n_e e \mathbf{v}_e$   $\rho \approx m_i n_i$ 

$$\mathbf{v} = (m_i \mathbf{v}_i + m_e \mathbf{v}_e) / \rho \approx \mathbf{v}_i \qquad p = p_i + p_e$$

• Continuity:

• Momentum:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \frac{d \mathbf{v}}{d t} = -\nabla p + \mathbf{J} \times \mathbf{B}$$

force balance -pdv/dt JxB



If  $d\mathbf{v}/dt = 0 \Rightarrow \mathbf{J} \times \mathbf{B} = \nabla p$ 

- Generalised Ohm's law:  $\mathbf{E} + \mathbf{v} \times \mathbf{B} = \eta \mathbf{J}$
- Maxwells equations, Adiabatic equation:

 $\frac{p}{\rho^{\gamma}} = \text{const.},$ 

### "MHD with anisotropy in velocity, pressure"

• Pressure different parallel and perpendicular to field due mainly to *directed* neutral beam injection



Momentum

$$\rho \left( \frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \cdot \mathbf{v} \right) = -\nabla \cdot \overline{\mathbf{P}} + \mathbf{J} \times \mathbf{B}$$





## **Helicon wave fields**

Chang, Hole, Caneses,, Chen, Blackwell, Corr



## Wave fields indicate anomalous resistivity

Chang, Hole, Caneses, Chen, Blackwell, Corr



 Add external antenna in dielectric tensor formulation and solve for wave fields (EMS code)

match to |B(z)| and ∠B(z) with enhancement in collision frequency to ~9.5 (ion-acoustic turbulence?)

[Chang et al Phys. Plas. 19, 083511 (2012)]

### Modulation in B introduces mode gaps Chang, Breizman, Hole

Helicon mode eigenfunction equation ⇔ continuous modes

$$k_z^2 \left( \frac{\partial}{\partial r} r \frac{\partial}{\partial r} r E_{\varphi} - m^2 E_{\varphi} \right) = -E_{\varphi} m \frac{\omega^2}{c^2} r \frac{\partial g}{\partial r} \qquad \Leftrightarrow \qquad \omega \sim \omega_c \frac{k_z^2}{\omega_r^2}$$

 Apply periodic modulation to axial velocity through axial field



modulation strength parameter



**3** 

 $\omega_p^2/\omega_c$ 



# Broken periodicity introduces gap modes

Chang, Breizman, Hole

Introduce defect into periodicity



[Chang, Breizman, Hole, Plasma Phys. Control. Fusion 55 (2013) 025003]

## Gap formation: a generic wave phenomena

Existence of frequency gaps generic wave phenomenon: (e.g. electron band gap in conductors, Bragg reflection in optical interference filters)



## Gap modes also present in Tokamaks

### • A zoo of gaps and gap modes



[Spectrum of n=3 resonant singular frequencies Heidbrink Phys. Plasmas, Vol. 9, No. 5, May 2002]

# Who cares?... the impact of gap modes

- Alfvén eigenmodes are driven by wave-particle resonance.
  - e.g. Wave-particle resonance

Surfer rides wave



Boat drives waves



 As energetic particles from beams, radio-frequency heated or fusion alphas collide with thermal population they slow and hit resonances



# Who cares?... the impact of gap modes

- Alfvén eigenmodes are driven by wave-particle resonance.
  - e.g. Wave-particle resonance

Surfer gets dumped



Boat drives waves



- As energetic particles from beams, radio-frequency heated or fusion alphas collide with thermal population they slow and hit resonances
- At large amplitude, Alfven eigenmodes can eject the driving particles from confinement, damaging wall and extinguishing collisional heating



## **KSTAR**

- KSTAR, a new ~\$300m superconducting long pulse tokamak
- MOU between ANU and KSTAR



## Suspected BAE mode activity in KSTAR

Hole, Ryu, Woo, Bak, Sharapov, Fitzgerald

- ➢ P<sub>NBI</sub> ~1.2MW,
- ➢ P<sub>ECRH</sub>∼ 200kW,
- > 80keV injection energy:
- >  $1/3 < v_{||} / v_A < 1, v_{||} = v \cdot B$ >  $210 < I_p < 407 \text{ kA},$ >  $3 < \langle n_e \rangle < 6 \times 10^{19} \text{ m}^{-3}.$

Although ions have  $v_{||} < v_A$ they can drive the mode through side band resonances  $v_A/3$ ,  $v_A/5$ ,  $v_A/7$ 







## Suspected BAE mode activity in KSTAR

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## **Global mode with reduced damping**

• Computed ion sound, shear Alfven wave continuum (CSCAS)



S

S

# **Beam changes threshold to BAE drive**



- Increasing core localisation  $\Rightarrow$  lowered instability threshold
- Mode driven by beams through  $v_A/3$  resonance

#### BAE modes in H-1 Bertram, Blackwell, Hole

• H-1 has similar modes, but more complex because fully 3D

![](_page_24_Figure_2.jpeg)

• CAS3D computed eigenmode for candidate BAE

![](_page_24_Figure_4.jpeg)

 Mode postulated to be driven by inverted (hollow) temperature profile

[Betram, Blackwell, Hole, Plasma Phys. Control. Fusion 54 (2012) 055009]

![](_page_25_Figure_0.jpeg)

# Expected impact of anisotropy

- Small angle  $\theta_{b}~$  between beam, field  $\Rightarrow$   $p_{\parallel}$  >  $p_{\perp}$  -
- Beam orthogonal to field,  $\theta_b = \pi/2 \Rightarrow p_\perp > p_\parallel$
- *p*<sub>||</sub> surfaces can be distorted and displaced inward relative to flux surfaces (i.e. Plasma)

[Cooper et al, Nuc. Fus. 20(8), 1980]

• If  $p_{\perp} > p_{\parallel}$ , an increase will occur in centrifugal shift :

[R. Iacono, A. Bondeson, F. Troyon, and R. Gruber, Phys. Fluids B 2 (8). August 1990]

![](_page_26_Figure_7.jpeg)

• Compute  $p_{\perp}$  and  $p_{||}$  from moments of distribution function, computed by Monte-Carlo collision code

[M J Hole, G von Nessi, M Fitzgerald, K G McClements, J Svensson, PPCF 53 (2011) 074021]

• Infer  $p_{\perp}$  from measured diamagnetic current  $J_{\perp}$  using  $\nabla p = J \times B$ [see V. Pustovitov, PPCF 52 065001, 2010 and references therein]

## **Equilibrium in toroidal symmetry**

 $\mathbf{J} \times \mathbf{B} = \nabla p \quad \begin{cases} \mathbf{B} \cdot \nabla p = 0 & \therefore \text{ No pressure gradient along B} \\ \mathbf{J} \cdot \nabla p = 0 & \therefore \text{ Current flows in magnetic surfaces} \end{cases}$ 

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J}, \qquad \nabla \cdot \mathbf{B} = 0$$

- Assume toroidal symmetry.
- Introduce

>co-ord. system ( $R, \zeta, z$ ) >poloidal flux function  $\psi$ ,  $\mathbf{B} \cdot \nabla \psi = 0$ >toroidal flux function  $f(\psi) = RB_{\zeta}(\psi, R)/\mu_0$ 

• Can reduce to Grad-Shafranov equation

$$\nabla \cdot \frac{1}{R^2} \nabla \psi = -\mu_0 \frac{J_{\varsigma}}{R} = -\mu_0 p'(\psi) - \frac{\mu_0^2}{R^2} f(\psi) f'(\psi)$$

• Normally solved by iterating calculation of field from current and current from field with prescribed boundary and  $\frac{\{p'(\psi), f(\psi)\}}{\{p'(\psi), f(\psi)\}}$ 

![](_page_27_Figure_9.jpeg)

# **Equilibrium with rotation & anisotropy**

• Inclusion of anisotropy and flow in equilibrium MHD equations

Momentum: 
$$\rho \mathbf{v} \cdot \nabla \mathbf{v} = \mathbf{J} \times \mathbf{B} - \nabla \cdot \overline{\mathbf{P}} \quad \overline{\mathbf{P}} = p_{\perp} \overline{\mathbf{I}} + \Delta \mathbf{B} \mathbf{B} / \mu_0, \quad \Delta = \frac{\mu_0 (p_{//} - p_{\perp})}{B^2}$$

• Frozen flux  $\Rightarrow$   $\mathbf{v} = -R\phi'_E(\psi)\mathbf{e}_{\varphi} = R\Omega(\psi)\mathbf{e}_{\varphi}$  Force balance becomes ...

$$\nabla \cdot \left[ (1 - \Delta) \left( \frac{\nabla \psi}{R^2} \right) \right] = -\frac{\partial p_{\parallel}}{\partial \psi} - \rho H'(\psi) + \rho \frac{\partial W}{\partial \psi} - \frac{f(\psi)f'(\psi)}{R^2(1 - \Delta)} + R^2 \rho \Omega(\psi) \Omega'(\psi)$$
Lorentz centrifugal
Bernoulli equation:
$$H(\psi) = W(\rho, B, \psi) - \frac{1}{2} \left[ R \phi'_E(\psi) \right]^2$$
Set of 5 profile constraints
$$\left\{ F(\psi), \Omega(\psi), H(\psi), \frac{\partial p_{\parallel}}{\partial \psi}, \frac{\partial W}{\partial \psi} \right\}$$

- $\partial W/ \; \partial \; \psi$  : different for MHD/ double-adiabatic/ guiding centre
- If two temperature Bi-Maxwellian model chosen

### New code written and benchmarked Fitzgerald, Hole, Appel

- New code EFIT TENSOR written to solve force balance with flow and anisotropy
- "Soloviev" analytic and real data benchmarks (MAST #13050, #18696) have been computed for isotropic, anisotropic and flow cases

![](_page_29_Figure_3.jpeg)

### Effect of anisotropy on MAST Magnetics Hole, Fitzgerald, von Nessi

- MAST #18696
- 1.9MW NB heating
- $I_p = 0.7 MA$ ,  $\beta_n = 2.5$
- Magnetics shows wave activity

![](_page_30_Figure_5.jpeg)

![](_page_30_Figure_6.jpeg)

[M.P. Gryaznevich et al, Nuc. Fus. 48, 084003, 2008]

- What is the impact on q profile due to presence of anisotropy and flow?
- How would this change wave activity?

# Monte Carlo collision code gives $p_{\parallel}$ , $p_{\perp}$ , v

![](_page_31_Figure_1.jpeg)

 $p_{\perp}/p_{\parallel} \approx 1.7$   $\rho = \sqrt{\Phi/\Phi_0} \approx s$   $\Phi = \text{toroidal flux}$  $M_{\phi,max} = 0.3$ 

Impact on plasma computed with EFIT TENSOR

![](_page_31_Figure_4.jpeg)

[M J Hole et al, PPCF 53 (2011) 074021]

# Impact of anisotropy on wave modes

15

10

-5

- 10

-15

-20

![](_page_32_Figure_1.jpeg)

- How do predicted mode frequencies change due to changes in q produced by anisotropy and flow?
- Calculation of change in stability due to anisotropy in progress...

*can address:* What is the change in ideal MHD modes of n=1 TAE?

Toroidal mode number identification

## Impact of anisotropy on wave modes

![](_page_33_Figure_1.jpeg)

[MJ Hole, G von Nessi, M Fitzgerald, PPCF 55 (2013) 014007]

# **Increased shear gives multiple TAEs**

![](_page_34_Figure_1.jpeg)

## **Increased shear gives multiple TAEs**

![](_page_35_Figure_1.jpeg)

## **Increased shear gives multiple TAEs**

![](_page_36_Figure_1.jpeg)

![](_page_37_Figure_0.jpeg)

# **Toroidal plasma equilibrium in 3D**

- In ideal MHD non-axisymmetric magnetic fields generally do not have a nested family of smooth flux surfaces, unless ideal surface currents are allowed at the rational surfaces.
- If the field is non-integrable (i.e. chaotic), then any continuous pressure that satisfies B·∇p=0 must have an infinitely discontinuous gradient, ∇p.
- Instead, solutions with stepped-pressure profiles are guaranteed to exist. Variational principle called MRXMHD (R. L. Dewar).
- Numerical implementation, SPEC, by S. Hudson (PPPL).

# **Taylor Relaxed States: Relaxed MHD**

- Zero pressure gradient regions are force-free magnetic fields:
- In 1974, Taylor argued that turbulent plasmas with small resistivity, and viscosity relax to a Beltrami field vacuum

nternal energy: 
$$W = \int_{P \cup V} \left( \frac{B^2}{2\mu_0} + \frac{p}{\gamma - 1} \right) d\tau^3$$
  
Total Helicity :  $H = \int_{V} (\mathbf{A} \cdot \mathbf{B}) d\tau^3$  interface I P

Taylor solved for minimum W subject to fixed H

i.e. solutions to  $\delta F=0$  of functional  $F = W - \mu H / 2$ 

$$P: \quad \nabla \times \mathbf{B} = \mu \mathbf{B}$$
  

$$I: \quad \left[ \left[ \frac{B^2}{2\mu_0} + p \right] \right] = 0$$
  

$$V: \quad \nabla \times \mathbf{B} = 0$$

# **Generalised Taylor Relaxation: MRXMHD**

R. L. Dewar, S. R. Hudson, M. J. Hole, G. Dennis, G. von Nessi

• Assume each invariant tori  $I_i$  act as ideal MHD barriers to relaxation, so that Taylor constraints are localized to subregions.

New system comprises:

- > N plasma regions  $P_i$  in relaxed states.
- > Regions separated by ideal MHD barrier  $I_i$ .
- > Enclosed by a vacuum V,
- Encased in a perfectly conducting wall W

$$W_l = \int_{R_l} \left( \frac{B_l^2}{2\mu_0} + \frac{P_l}{\gamma - 1} \right) d\tau^3 \qquad P_l:$$

$$H_l = \int_V (\mathbf{A}_l \cdot \mathbf{B}_l) d\tau^3$$

Seek minimum energy state:

$$F = \sum_{l=1}^{N} \left( W_l - \mu_l H_l / 2 \right)$$

![](_page_40_Picture_12.jpeg)

$P_l$ :	$ abla  imes {f B} = \mu_l {f B}$
	$P_l = \text{constant}$
$I_l$ :	$\mathbf{B} \cdot \mathbf{n} = 0$
	$[[P_l + B^2 / (2\mu_0)]] = 0$
V :	$\nabla \times \mathbf{B} = 0$
	$\nabla \cdot \mathbf{B} = 0$
W:	$\mathbf{B} \cdot \mathbf{n} = 0$

![](_page_41_Figure_0.jpeg)

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_42_Figure_3.jpeg)

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_43_Figure_3.jpeg)

![](_page_43_Figure_4.jpeg)

• Model RFX-mod QSH state by a 2-interface minimum energy MRXMHD state.

[G. R. Dennis submitted Phys. Rev. Lett. 22/02/2013 http://arxiv.org/abs/1302.5458]

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_44_Figure_3.jpeg)

Dennis, Hudson, Terranova, Dewar, Hole

 The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_45_Figure_3.jpeg)

"Experimental" Poincaré plot

[Fig. 6 of P. Martin et al., Nuclear Fusion 49, 104019 (2009)]

[G. R. Dennis submitted Phys. Rev. Lett. 22/02/2013 http://arxiv.org/abs/1302.5458]

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_46_Figure_3.jpeg)

"Experimental" Poincaré plot

[Fig. 6 of P. Martin et al., Nuclear Fusion 49, 104019 (2009)]

- Model RFX-mod QSH state by a 2-interface minimum energy MRXMHD state.
- Theoretical Poincare plots match experiment as barrier is changed

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_47_Figure_3.jpeg)

"Experimental" Poincaré plot

[Fig. 6 of P. Martin et al., Nuclear Fusion 49, 104019 (2009)]

- Model RFX-mod QSH state by a 2-interface minimum energy MRXMHD state.
- Theoretical Poincare plots match experiment as barrier is changed

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_48_Figure_3.jpeg)

"Experimental" Poincaré plot

[Fig. 6 of P. Martin et al., Nuclear Fusion 49, 104019 (2009)]

![](_page_48_Picture_6.jpeg)

- Model RFX-mod QSH state by a 2-interface minimum energy MRXMHD state.
- Theoretical Poincare plots match experiment as barrier is changed
- RFP bifurcated state has lower energy (preferred) than comparable axis-symmetric state

Dennis, Hudson, Terranova, Dewar, Hole

• The quasi-single helicity state is a stable helical state in RFP: becomes purer as current is increase

![](_page_49_Figure_3.jpeg)

"Experimental" Poincaré plot

[Fig. 6 of P. Martin et al., Nuclear Fusion 49, 104019 (2009)]

![](_page_49_Picture_6.jpeg)

- Model RFX-mod QSH state by a 2-interface minimum energy MRXMHD state.
- Theoretical Poincare plots match experiment as barrier is changed
- RFP bifurcated state has lower energy (preferred) than comparable axis-symmetric state
- Looks similar to H-1 (next step), indeed ... next stellarator conference is joint with RFP's

![](_page_50_Figure_0.jpeg)

### Bayesian equilibrium modelling: integrating data with theory J. Svensson, G. von Nessi, M. Hole, L. Appel,

 $P(\mathbf{H}|\mathbf{D}) = P(\mathbf{D}|\mathbf{H})P(\mathbf{H})/P(\mathbf{D})$ 

 $\mathbf{H} = \left\{ J_{\phi}(R,Z), p'(\psi), f(\psi), \rho(\psi,R), \Omega(\psi) \right\}$ 

 $\mathbf{D} = \left\{ P_i(R, Z), F_i(R, Z), \text{ tan } \gamma_i(R, Z), I_p, P_{s,e}, S_e(k, \omega), S_C(\nu) \right\}$ 

Pick-up Flux MSE Plasma TS CXRS coils loops signals current spectra spectra

(1) Improve equilibrium reconstruction

(2) Validate different physics models

Two fluid with rotation

Aims

[McClements & Thyagaraja Mon. Not. R. Astron. Soc. 323 733-42 2001]

Ideal MHD fluid with rotation

[Guazzotto L et al, Phys. Plasmas 11 604-14, 2004]

Energetic particle resolved multiple-fluid

[Hole & Dennis, PPCF 51 035014, 2009]

(3) Infer poorly diagnosed physics parameters

![](_page_52_Figure_0.jpeg)

## **ITER:** The next step for fusion power

![](_page_53_Picture_1.jpeg)

- Fusion power = 500MW
- Power Gain > 10
- Temperature ~ 100 million °C
- Growing Consortium

![](_page_53_Figure_6.jpeg)

Collaboration agreements with

International Atomic Energy Agency
 CERN – world's largest accelerator
 Principality of Monaco

Construction +10 year operation cost ~\$20 billion *Fiscally, world's largest science experiment* 

## **Plasma Physics Challenges**

- Production and study of a plasma dominated by self heating. *Burning Plasma Physics.*
- New instabilities in burning plasmas: possibilities energetic particle modes driven by beam ions, fusion  $\alpha$ s could "short-circuit" heating of thermal plasma
- Edge Localised Modes: Control of field lines that erupt through plasma edge
- Better disruption mitigation (e.g. massive gas puff injection).
- Real time mode control and identification
- Measurement and "integrated modelling" of plasmas under extreme conditions.

![](_page_54_Picture_7.jpeg)

Difference images from  $D^{\alpha}$  camera of MAST plasmas

# **Physics contributions to ITER**

	equilibriu	n waves	stability	wave-wave	wave-plasma nonlinearity		
					stochastic		
		gap mode physics		Continuum damping			
mplex	ITER						
	Bayesian	Anisotropy					
gurati	3D fields: ripple,						
confi	error, RMP						

# Australian contributions to ITER science Powering ahead:

a national response to the rise of the international fusion power program

- B. James (coordinator), R. Garrett, B. Green, M. J. Hole, J. Howard, J. O'Connor
- Successor to a strategy for fusion science in Australia released in 2007
- Outcomes: H-1 upgrade, influence Future Fellowship Scheme

Exposure draft available at

http://www.ainse.edu.au/fusion/iter/australian\_fusion\_strategy2

- Proposed participation in the International Tokamak Physics Activity, coordinated by the ITER Organisation. ITPA topical groups in
  Diagnostics
  Energetic Particle Physics
  Integrated Operating Scenarios
  MHD, Disruption and Control
  Pedestal and Edge
  Scrape-off Layer and Divertor
  - Transport and Confinement

- Construction of a pilot coherence imaging divertor diagnostic
- Country level agreement proposed through ANSTO

![](_page_57_Picture_0.jpeg)

- Plasma Theory and Modelling: a vibrant ANU pursuit developing theory for next generation fusion experiments, and supporting physics interpretation of existing experiments.
- Very strong international collaboration
- Research areas
  - Burning plasma physics: anisotropy and flow, energetic particle driven modes
  - > 3D MHD physics (e.g. MRXMHD). Impacts of 3D structure on plasma.
  - Bayesian inference of configurations
  - Interpretation/modelling of international and domestic experiments
  - > Not mentioned... continuum damping, pulsar modelling, ELM statistics
- Research synergies identified with ITER and its strategic research community, the International Tokamak Physics Activity ITPA. These align with Australian strategic planning for fusion science.